Gravitational Fields Question paper 4

| Level | International A Level | | |
|------------|-----------------------|--|--|
| Subject | Physics | | |
| Exam Board | CIE | | |
| Торіс | Gravitational Fields | | |
| Sub Topic | | | |
| Paper Type | Theory | | |
| Booklet | Question paper 4 | | |

| Time Allowed: | 64 minutes | | |
|---------------|------------|--|--|
| Score: | /53 | | |
| Percentage: | /100 | | |

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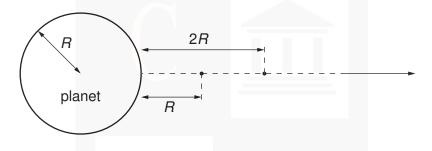
| A* | A | В | С | D | E | U |
|------|--------|-----|-------|-------|-----|------|
| >85% | '77.5% | 70% | 62.5% | 57.5% | 45% | <45% |

1 (a) Explain what is meant by a *gravitational field*.

......[1]

(b) A spherical planet has mass M and radius R. The planet may be considered to have all its mass concentrated at its centre.

A rocket is launched from the surface of the planet such that the rocket moves radially away from the planet. The rocket engines are stopped when the rocket is at a height R above the surface of the planet, as shown in Fig. 1.1.





The mass of the rocket, after its engines have been stopped, is m.

(i) Show that, for the rocket to travel from a height *R* to a height 2*R* above the planet's surface, the change $\Delta E_{\rm P}$ in the magnitude of the gravitational potential energy of the rocket is given by the expression

$$\Delta E_{\rm P} = \frac{GMm}{6R}.$$

(ii) During the ascent from a height *R* to a height 2*R*, the speed of the rocket changes from 7600 m s⁻¹ to 7320 m s⁻¹. Show that, in SI units, the change $\Delta E_{\rm K}$ in the kinetic energy of the rocket is given by the expression

$$\Delta E_{\rm K} = (2.09 \ 10^6)m.$$



[1]

- (c) The planet has a radius of $3.40 \quad 10^6$ m.
 - (i) Use the expressions in (b) to determine a value for the mass *M* of the planet.

..... kg [2] M (ii) State one assumption made in the determination in (i).[1]

2 A rocket is launched from the surface of the Earth.

Fig. 4.1 gives data for the speed of the rocket at two heights above the Earth's surface, after the rocket engine has been switched off.

| height / m | speed / m s ⁻¹ |
|-----------------------|---------------------------|
| $h_1=19.9\times 10^6$ | v ₁ = 5370 |
| $h_2=22.7\times 10^6$ | v ₂ = 5090 |



The Earth may be assumed to be a uniform sphere of radius $R = 6.38 \times 10^6$ m, with its mass *M* concentrated at its centre. The rocket, after the engine has been switched off, has mass *m*.

- (a) Write down an expression in terms of
 - (i) G, M, m, h_1 , h_2 and R for the change in gravitational potential energy of the rocket,

......[1]

(ii) m, v_1 and v_2 for the change in kinetic energy of the rocket.

......[1]

(b) Using the expressions in (a), determine a value for the mass *M* of the Earth.

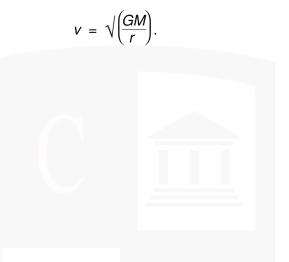


M = kg [3]

3 The Earth may be considered to be a uniform sphere with its mass *M* concentrated at its centre.

A satellite of mass *m* orbits the Earth such that the radius of the circular orbit is *r*.

(a) Show that the linear speed v of the satellite is given by the expression



- [2]
- (b) For this satellite, write down expressions, in terms of G, M, m and r, for
 - (i) its kinetic energy,

kinetic energy =[1]

(ii) its gravitational potential energy,

potential energy =[1]

(iii) its total energy.

(c) The total energy of the satellite gradually decreases.

State and explain the effect of this decrease on

(i) the radius *r* of the orbit,
(ii) the linear speed *v* of the satellite.
[2]



- 4 The Earth may be considered to be a sphere of radius 6.4×10^6 m with its mass of 6.0×10^{24} kg concentrated at its centre. A satellite of mass 650 kg is to be launched from the Equator and put into geostationary orbit.
 - (a) Show that the radius of the geostationary orbit is 4.2×10^7 m.



[3]

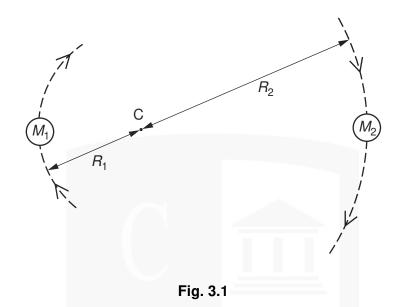
(b) Determine the increase in gravitational potential energy of the satellite during its launch from the Earth's surface to the geostationary orbit.

energy = J [4]

(c) Suggest one advantage of launching satellites from the Equator in the direction of rotation of the Earth.

.....[1]

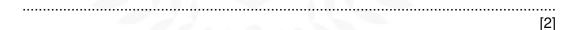
5 A binary star consists of two stars that orbit about a fixed point C, as shown in Fig. 3.1.



The star of mass M_1 has a circular orbit of radius R_1 and the star of mass M_2 has a circular orbit of radius R_2 . Both stars have the same angular speed ω , about C.

- (a) State the formula, in terms of G, M_1 , M_2 , R_1 , R_2 and ω for
 - (i) the gravitational force between the two stars,

(ii) the centripetal force on the star of mass M_1 .



(b) The stars orbit each other in a time of 1.26×10^8 s (4.0 years). Calculate the angular speed ω for each star.

angular speed = $rad s^{-1}$ [2]

(c) (i) Show that the ratio of the masses of the stars is given by the expression

$$\frac{M_1}{M_2} = \frac{R_2}{R_1}.$$

[2]

(ii) The ratio $\frac{M_1}{M_2}$ is equal to 3.0 and the separation of the stars is 3.2×10^{11} m. Calculate the radii R_1 and R_2 .

1 = m

- $R_2 = \dots m_1$
 - [2]
- (d) (i) By equating the expressions you have given in (a) and using the data calculated in (b) and (c), determine the mass of one of the stars.

mass of star = kg

(ii) State whether the answer in (i) is for the more massive or for the less massive star.

6 (a) (i) On Fig. 1.1, draw lines to represent the gravitational field outside an isolated uniform sphere.



Fig. 1.1

(ii) A second sphere has the same mass but a smaller radius. Suggest what difference, if any, there is between the patterns of field lines for the two spheres.



(b) The Earth may be considered to be a uniform sphere of radius 6380 km with its mass of $5.98 \times 10^{24} \text{ kg}$ concentrated at its centre, as illustrated in Fig. 1.2.

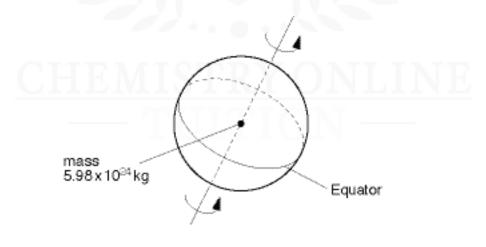


Fig. 1.2

A mass of 1.00 kg on the Equator rotates about the axis of the Earth with a period of 1.00 day $(8.64 \times 10^4 s)$.

Calculate, to three significant figures,

(i) the gravitational force $F_{\rm G}$ of attraction between the mass and the Earth,

